

LOFAR surveys: The formation and evolution of galaxies, AGN and clusters of galaxies

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Members Core team Röttgering¹²³⁴⁹(Leiden), Barthel³⁴⁵⁶(Groningen), Best¹²³⁴⁵ (Edinburgh), Brüggen² (Bremen), Brunetti² (Bologna), Chyzy²³⁶ (Kraków), Conway^{56t} (Göteborg), Jarvis¹³⁴⁹ (Hertfordshire), Lehnert³⁶ (Meudon), Miley¹²⁴⁵ (Leiden), Morganti⁴⁵ (Dwingeloo), Wise²⁴⁵ (ASTRON)

Regular members: Haverkorn⁸ (ASTRON), Jackson⁷ (Manchester), White³⁸ (Open University), Abdalla⁹ (UCL London), Anderson (MPIfR Bonn), Arnaud² (Meudon), Bacon⁷⁹ (Portsmouth), Beck⁶ (Bonn), Beswick⁵⁶ (Manchester), Brentjens² (ASTRON), Britzen⁵ (Bonn), Conselice (Nottingham), Croston² (Southampton), Dettmar⁶ (Bochum), Eales⁶ (Cardiff), Edge² (Durham), Engels⁴ (Hamburg), Enßlin² (Garching), Falcke¹⁴⁵ (Nijmegen), Feretti² (Bologna), Ferrari² (Nice), Franx³ (Leiden), Garrett³⁷ (ASTRON), Génova-Santos¹ (IAC), Hardcastle (Hertfordshire), Hendry⁹ (Glasgow), Hoeft² (Tautenburg), Horellou²⁵⁶ (Onsala), Isral⁶ (Leiden), Ivison³ (Edinburgh), Jamroz⁴⁵ (Krakow), Kassim⁸ (Washington), Kauffmann⁴ (Garching), Klein⁶ (Bonn), Kuijken⁷ (Leiden), Kunert-Bajraszewska⁴⁵ (Torun), Lobanov⁵ (Bonn), Marecki (Torun), Marti-Vidal⁶ (Onsala), Martinez-Sansigre (Portsmouth), McKean¹⁷ (ASTRON), Merloni⁴⁵ (Garching), Middelberg⁴ (Bochum), Murgia⁴⁵ (IAC), Nichol⁹ (Portsmouth), Oliver³ (Sussex), Oosterloo⁶ (ASTRON), Otmianowska-Mazur (Krakow), Page⁴ (London), Paragi (JIVE), Pentericci¹³ (Rome), Percival⁹ (Portsmouth), Peters⁸ (Washington), Polatidis⁵ (ASTRON), Prandoni³⁴ (IAC), Raychaudhury² (Birmingham), Reich⁸ (Bonn), Schwarz⁹ (Bielefeld), Simpson¹⁴ (Liverpool), Steinmetz⁵ (Potsdam), Strom⁵⁶⁸ (ASTRON), Tadhunter⁵ (Sheffield), Valentijn²⁶ (Groningen), van der Werf³ (Leiden), van Driel⁶ (Meudon), van Weeren¹²⁸ (ASTRON/Leiden), Varenius⁶ (Gothenburg), Vink⁸ (Amsterdam), White⁴ (Garching), Wisotzki⁴ (Potsdam), Wucknitz⁷ (Bonn), Zarb-Adami⁹ (Oxford), Zensus⁵ (Bonn)

Postdocs: Asgekar⁸ (ASTRON), Bertacca⁹ (UWC), Birzan²³⁵ (Leiden), Bonafede² (Bremen), Bonfield⁹ (Hertfordshire), Cassano² (IAC), Deller (ASTRON), Dwelly³ (Southampton), Faltenbacher⁹ (UWC), Heald⁶ (ASTRON), Heesen⁵⁶ (Hertfordshire), Heywood⁹ (Oxford), Johnston⁹ (UWC), Kapinska (Portsmouth), Kloeckner³⁴ (Oxford), König (Köln), Macario² (Nice), Mahony (ASTRON), Mauch³⁴ (Oxford), McKay (Chilbolton), McKee¹ (Leiden), Oonk⁸ (ASTRON), Orru¹²³⁵ (Nijmegen), Patel⁹ (Portsmouth), Pizzo² (ASTRON), Raccanelli⁹ (Portsmouth), Rafferty²³⁵ (Leiden), Sabater Montes⁴ (Edinburgh), Seymour¹ (Sydney), Smith⁹ (Herts), Smith⁹ (UWC), Stewart (Bonn), Tasse⁴ (Meudon), Tudose (ASTRON), Vaccari⁹ (UWC), van Bemmelen (ASTRON), Zwart⁹ (UWC)

PhDs: Batejat⁵⁶ (Gothenburg), De Gasperin⁴⁵ (Garching), Deane³ (Oxford), Drzazga³⁶ (Krakow), Fielding⁴ (Edinburgh), Guglielmino⁴⁵ (Bologna), Harwood⁵ (Hertfordshire), Heidenreich² (Southampton), Israel³ (Leiden), Junkelwitz² (Garching), Jurusik⁶ (Krakow), Ker¹³⁴ (Edinburgh), Kuligowska⁴⁵ (Krakow), Lazell² (Birmingham), Lindsay⁹ (Hertfordshire), Madhanpall⁹ (UWC), McAlpine⁹ (UWC), Morabito¹ (Leiden), Natt⁸ (Open University), Ogren² (Bremen), Rubart⁹ (Bielefeld), Shulevski⁵ (Groningen), Stroe² (Leiden), Temourian¹ (Hertfordshire), Trasatti² (Bonn), van Velzen¹ (Nijmegen), Williams⁴⁶ (Leiden).

Science working groups with chairs:

- | | |
|---|--|
| 1. High redshift radio galaxies - Miley | 5. Physics of nearby AGN - Morganti |
| 2. Galaxy clusters - Brüggen/ Brunetti | 6. Nearby galaxies - Conway/ Chyzy |
| 3. Cosmic Star-Formation - Lehnert/ Barthel | 7. Strong lensing - Jackson |
| 4. AGN and black hole evolution - Best | 8. The Galactic plane - Haverkorn/ White |
| | 9. Cosmology - Jarvis/ Bacon |

LOFAR
Survey team
PI HR

International LOFAR Telescope (ILT)



Chilbolton



Dutch stations



Onsala

Norderstedt

LOFAR Core (NL)

Potsdam

Baldy

Borówiec

Łazy

Tautenburg

Jülich

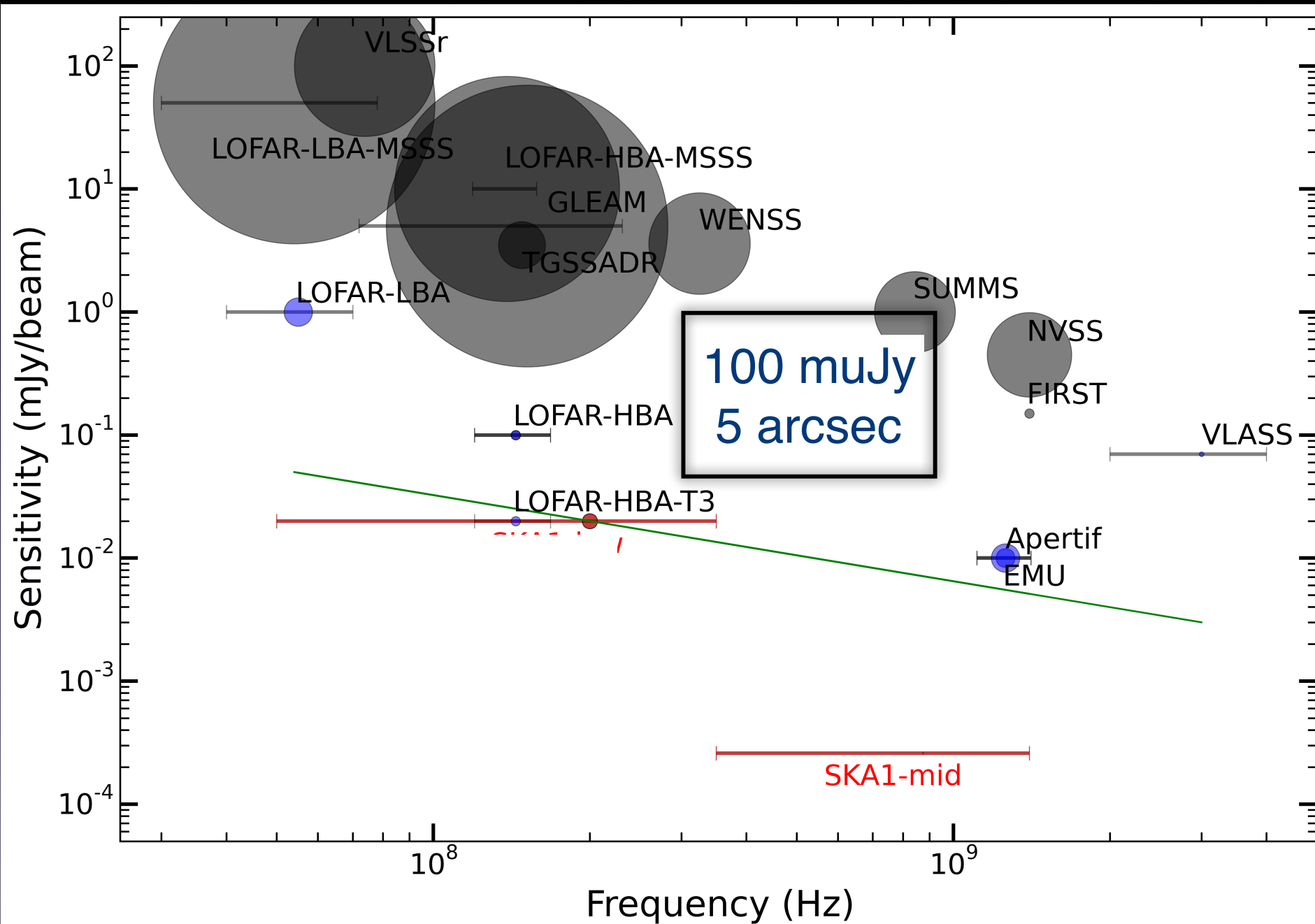
Effelsberg

Unterweilenbach

Nançay



- 47 operational stations completed
- 38 NL stations, 9 international stations
- 3 new stations coming in Poland



- 12 % of the northern sky already mapped

Activities

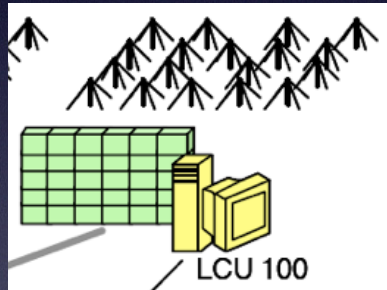
- Survey planning and production (Shimwell)
- Techniques, also to learn how to deal with SKA
 - Calibration techniques (De Gasperin, Intema, Dijkema)
 - Ionospheric simulation (Albert, Mevius)
 - Software profiling, grid enabling, speed and memory optimisation (Mechev)
 - Identifications, phot-z and id characterisations (Duncan)

Science

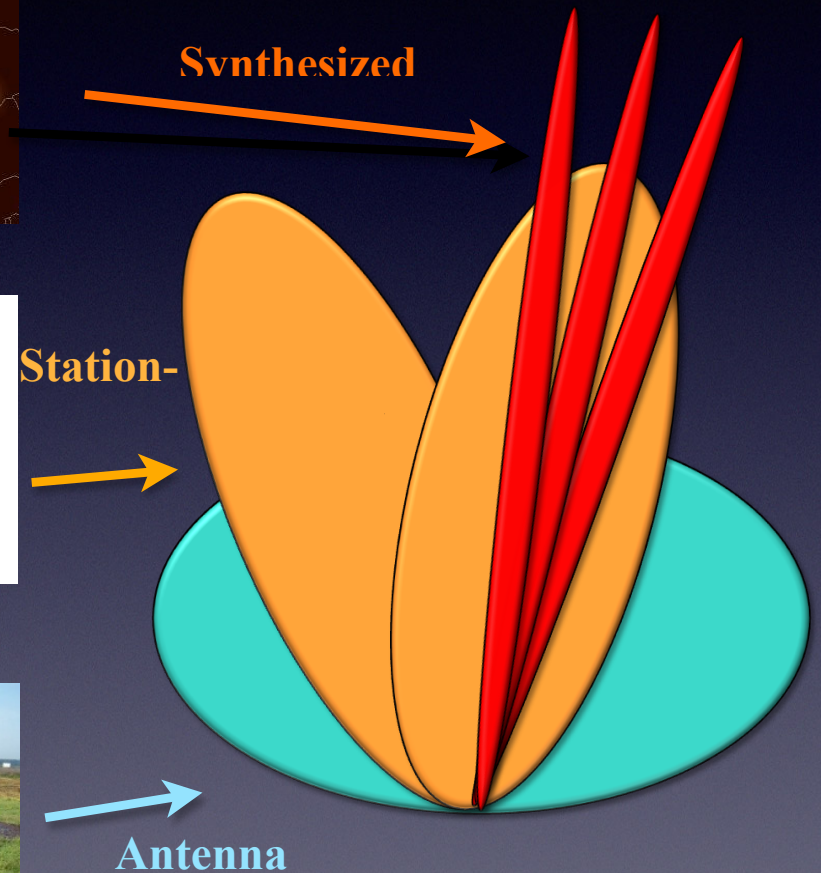
- The most distant radio galaxies to probe the EoR (Miley, Saxena, Retana, Morabito, Stroe, Falcke)
- Carbon recombination lines (Tielens, Emig, Morabito, Oonk, Sales-Munoz)
- Starburst galaxies and proto-clusters (Duncan, Rivera)
- Galaxy (Vink, Haverkorn)
- AGN feedback and the characterisation of AGN hosts (Jansen, Rivera, Williams, Morganti, McKean)
- Merging clusters (Hoang, Shimwell, Intema, de Gasparin, Mandal, Donnert, Albert, Wise)

Challenges

- Data rates of up to a Tb/s
- Radio interference
- Demixing: removal of A-team sources*
- Timing of the stations clocks
- Ionosphere
- Beam calibration
- Wide-field mapmaking

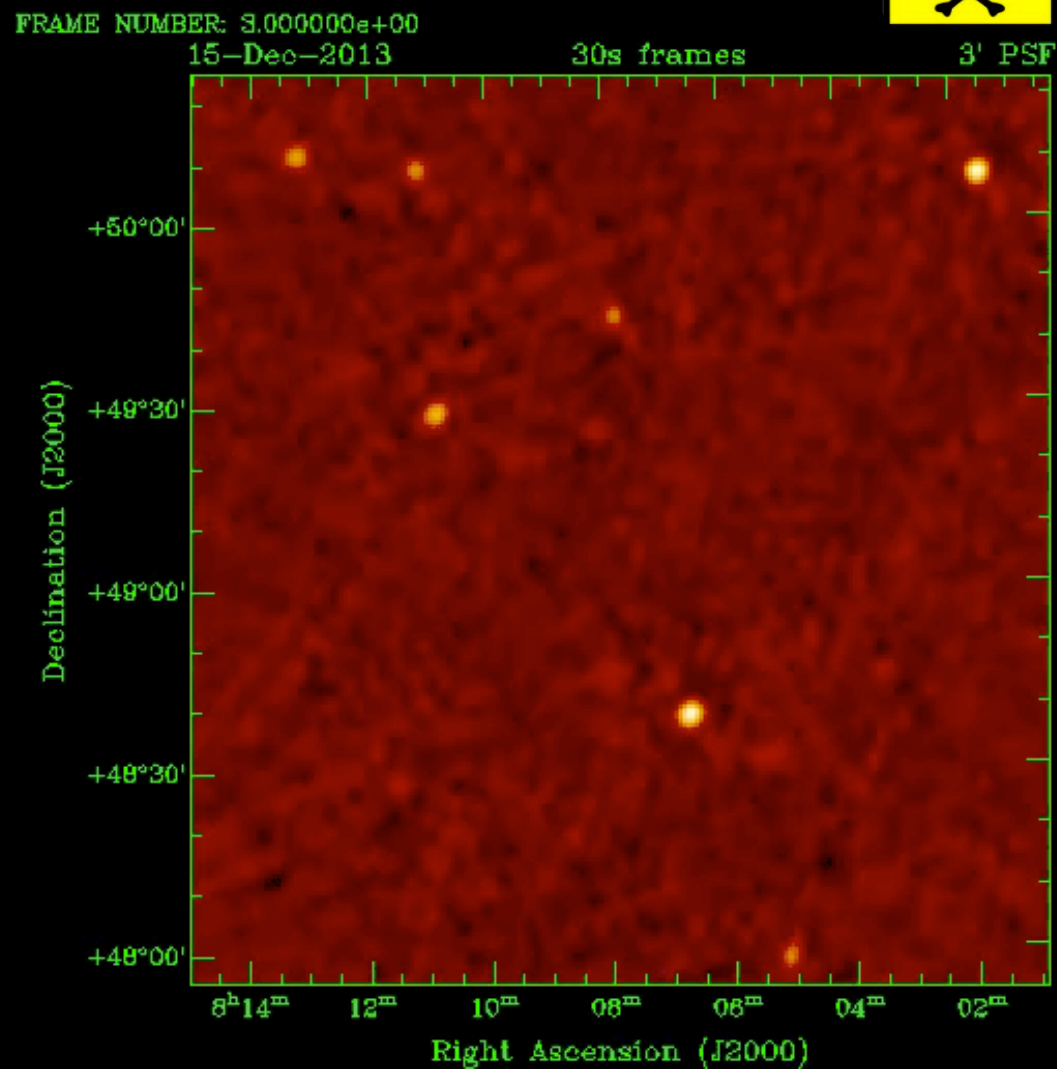
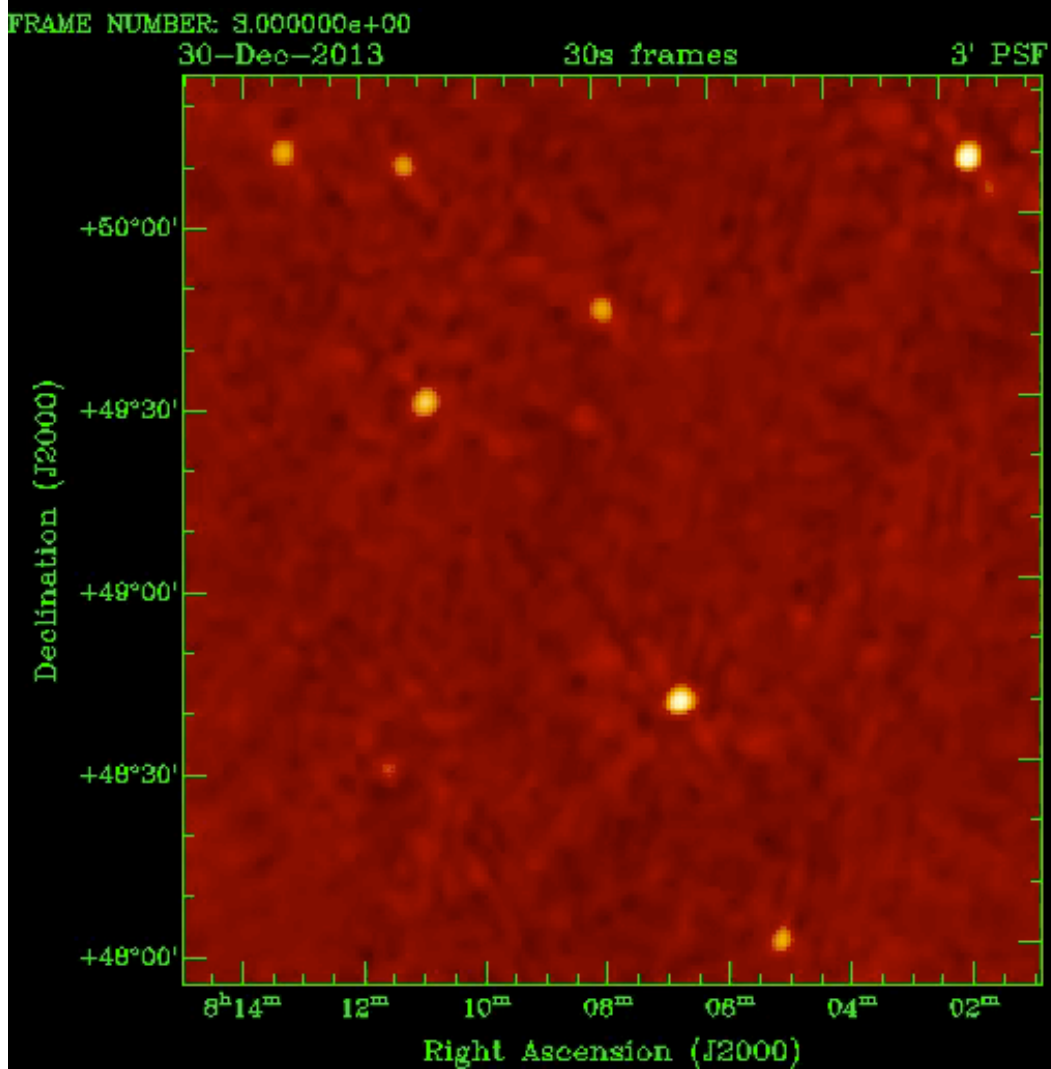


Three Beams

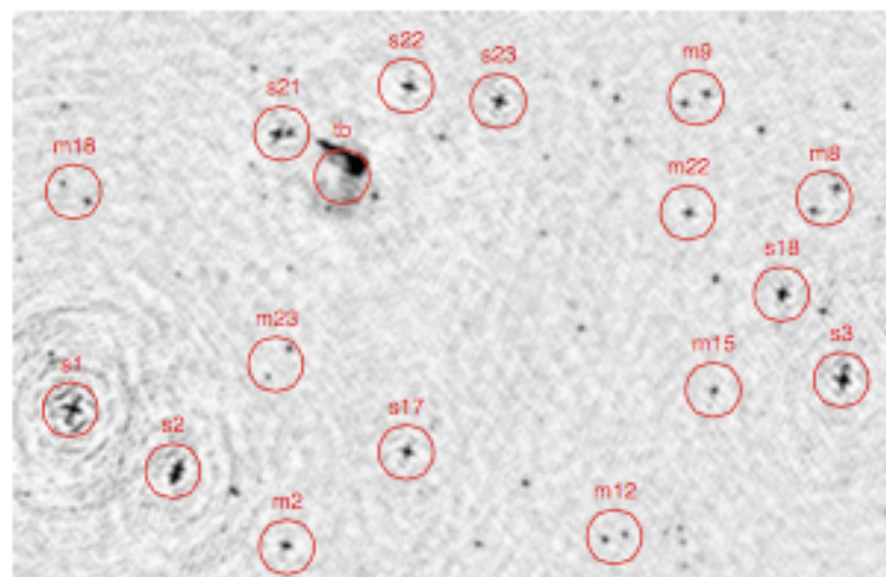
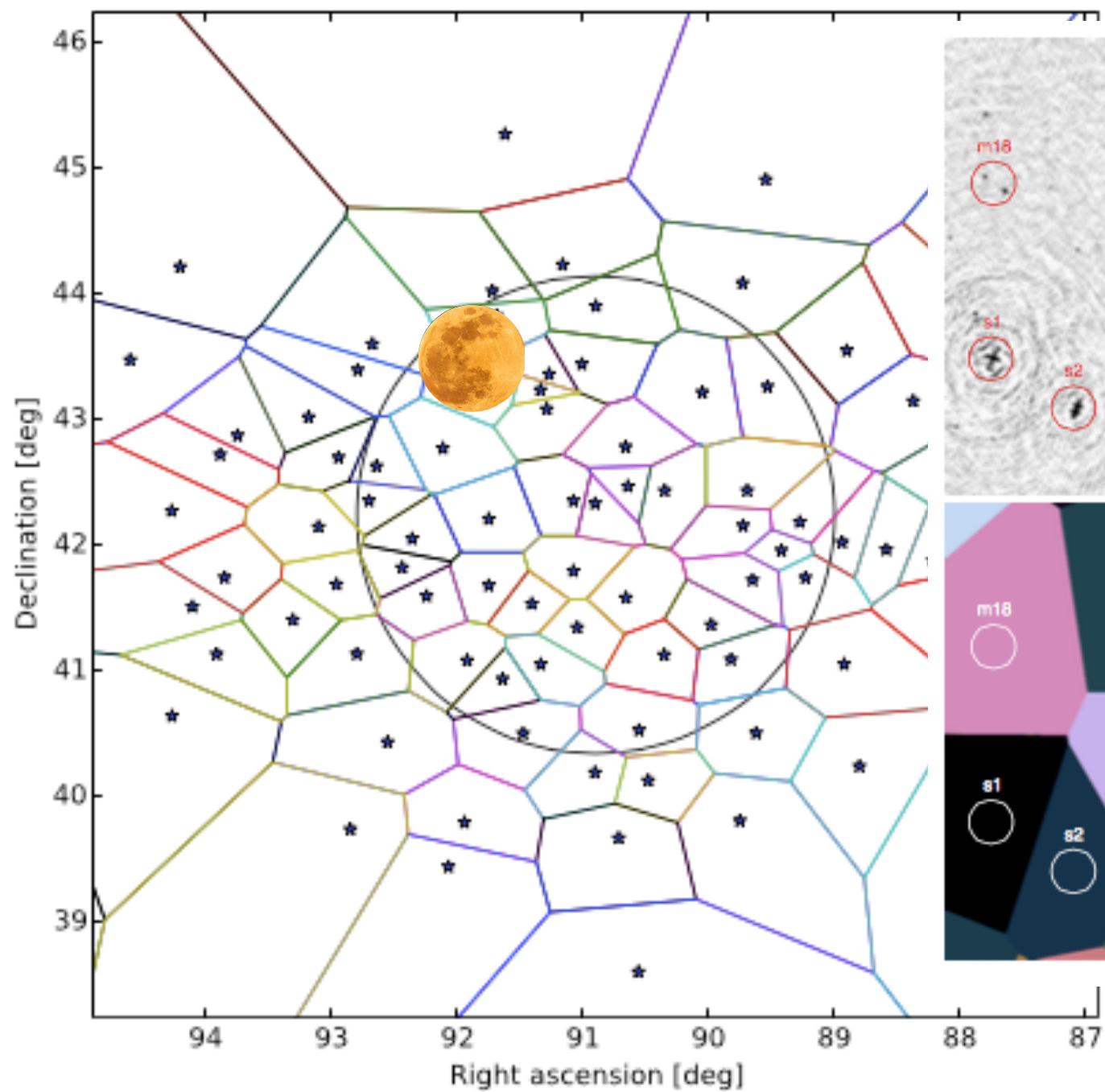


quiet ionosphere

wild ionosphere



Note: these images have only 3 arcmin resolution, the NL array has 5 arcsec resolution...

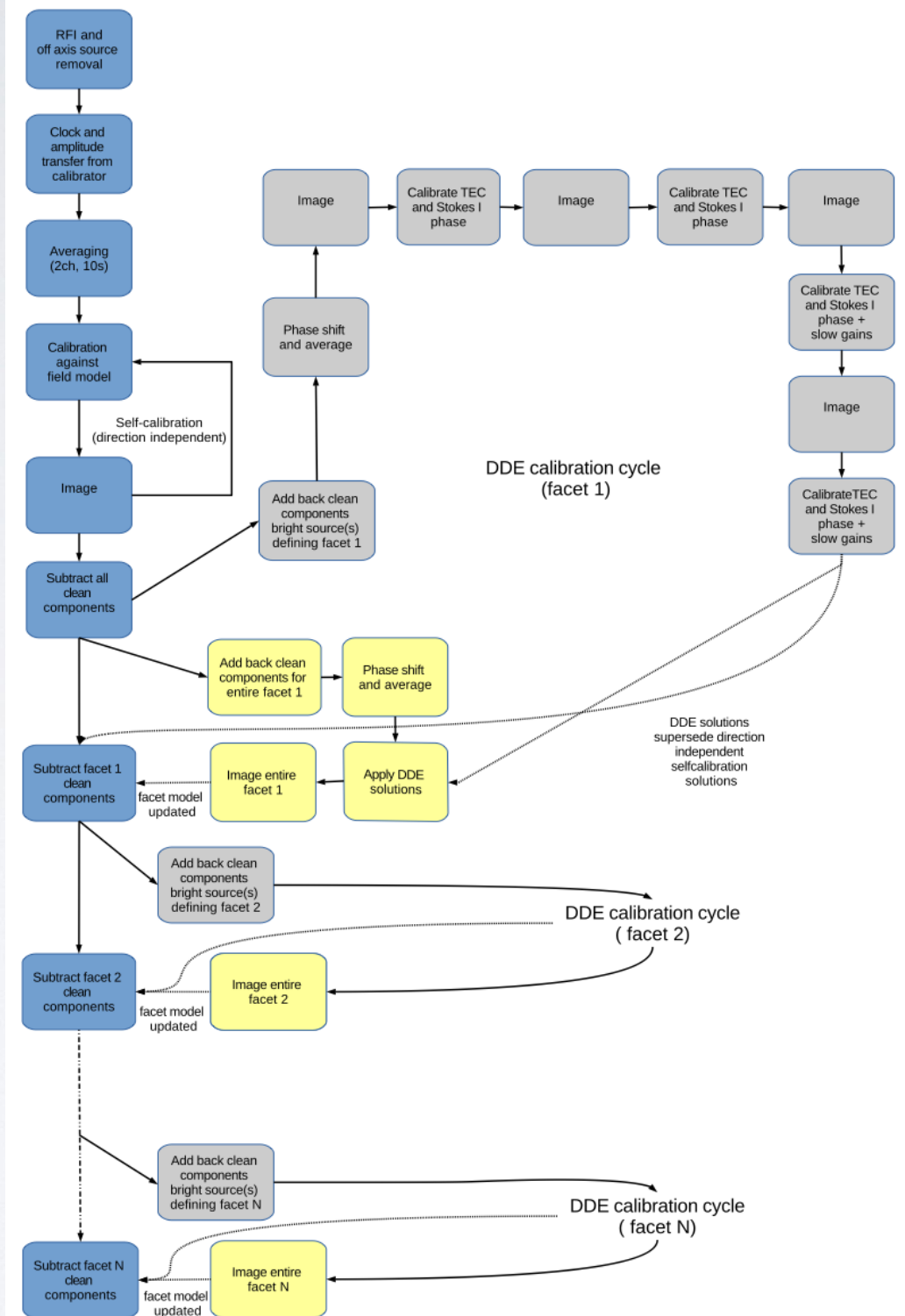


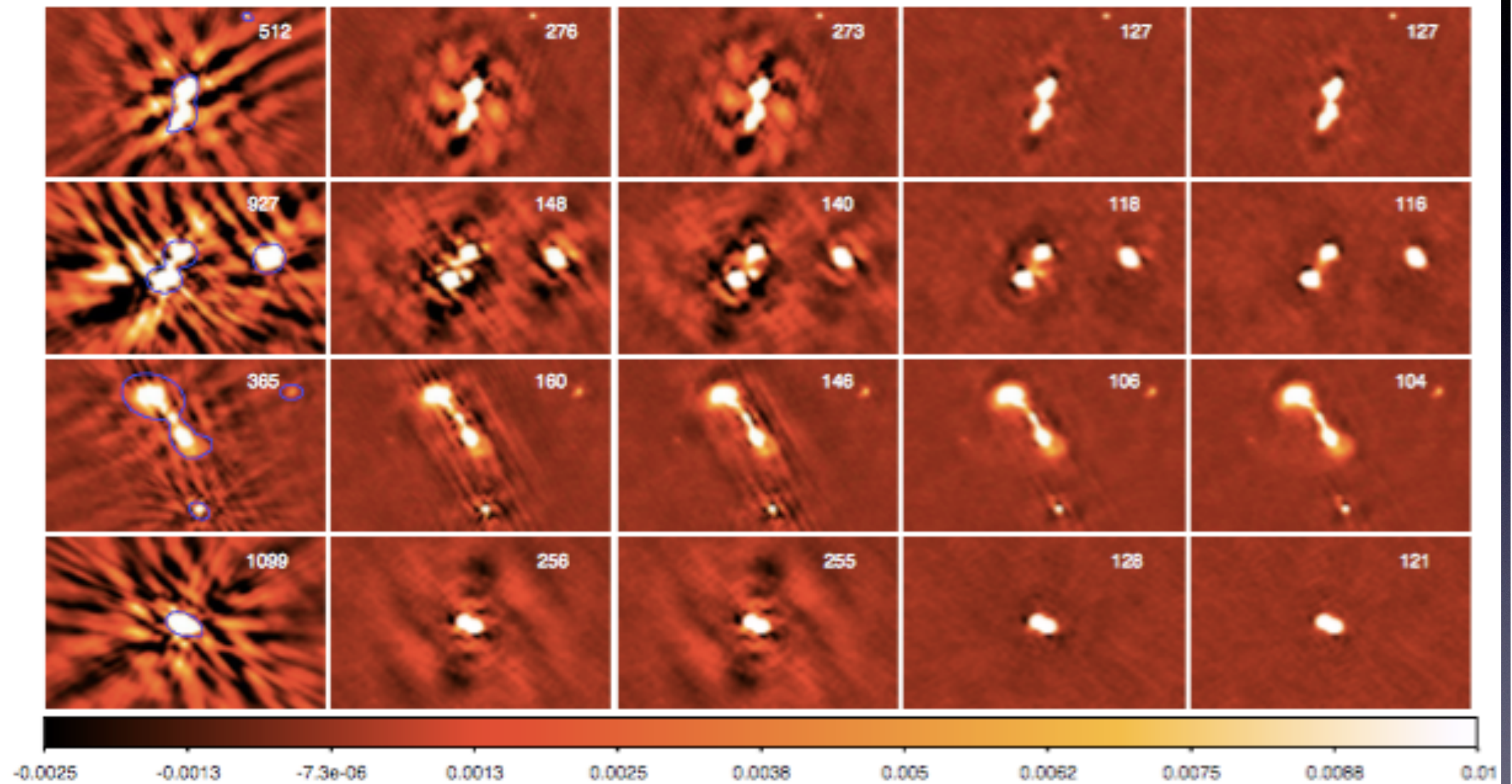
FACET CALIBRATION

- Subtract all sources from data
- Define facet centers

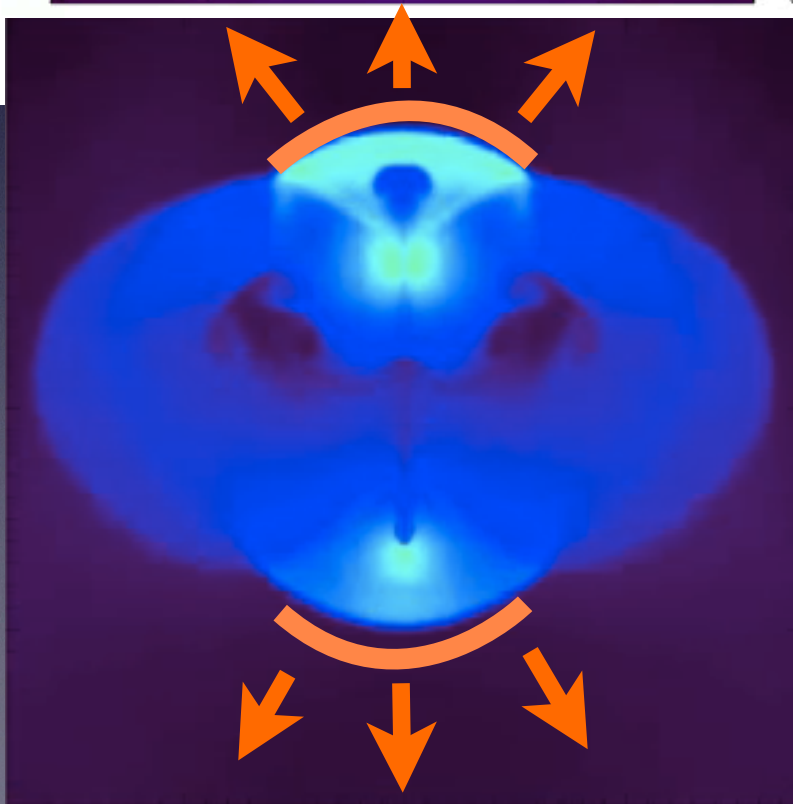
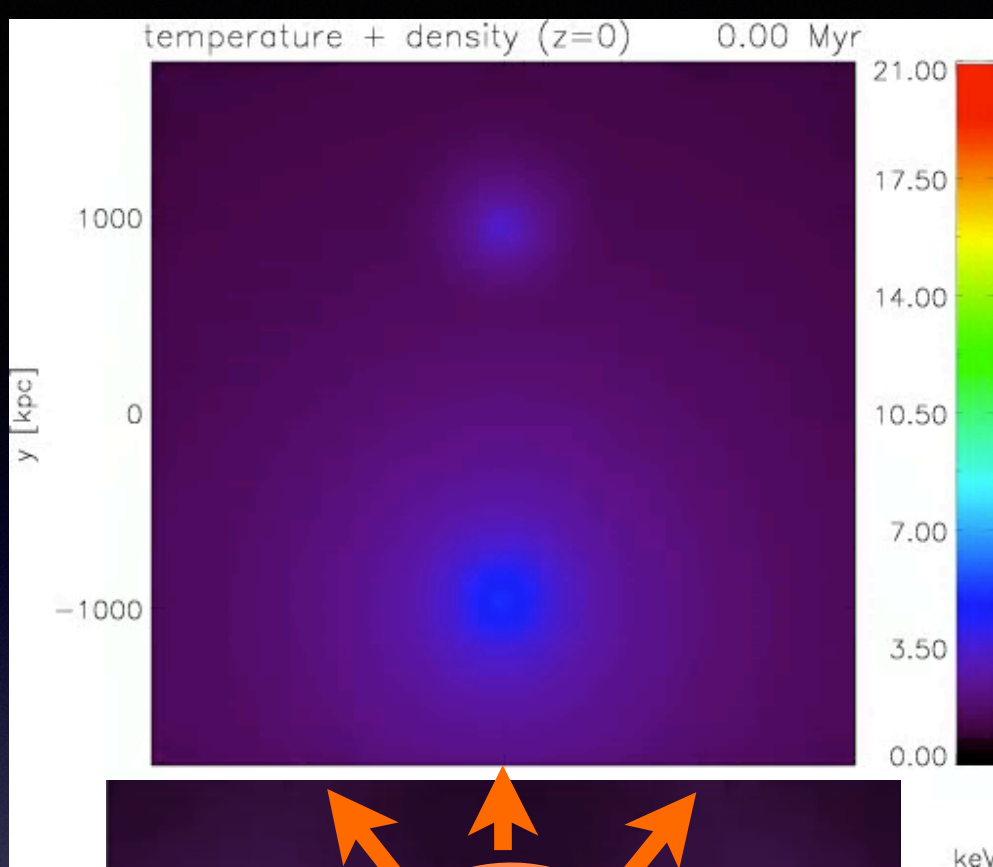
- Add back central source(s) defining facet
- Phase shift + average
- *DDE self-calibration Cycle*
- Add back all sources in facet
- Correct with solutions
- Image
- Subtract updated facet model with solutions

loop over facet



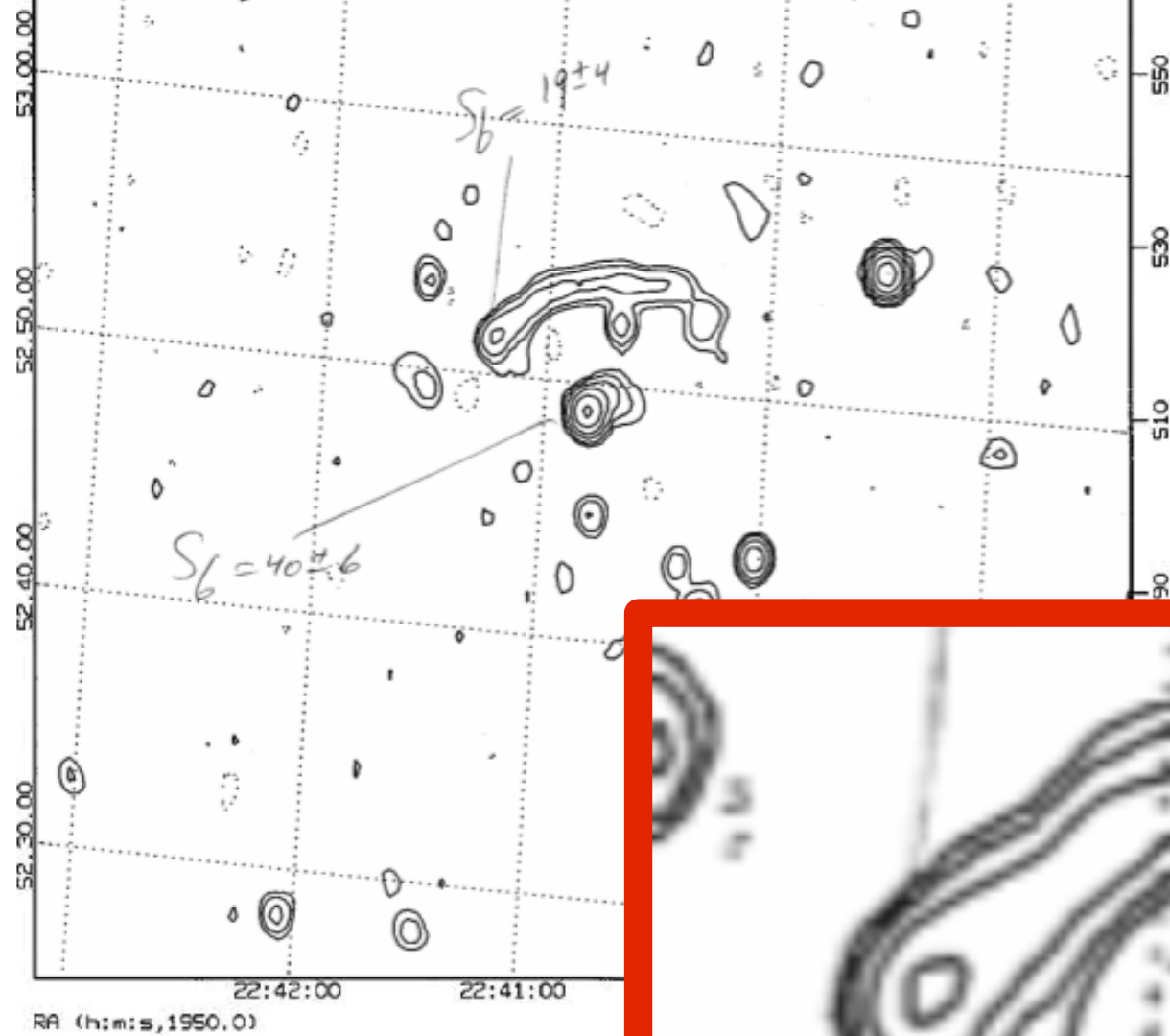


Simulation X-ray gas of two colliding clusters



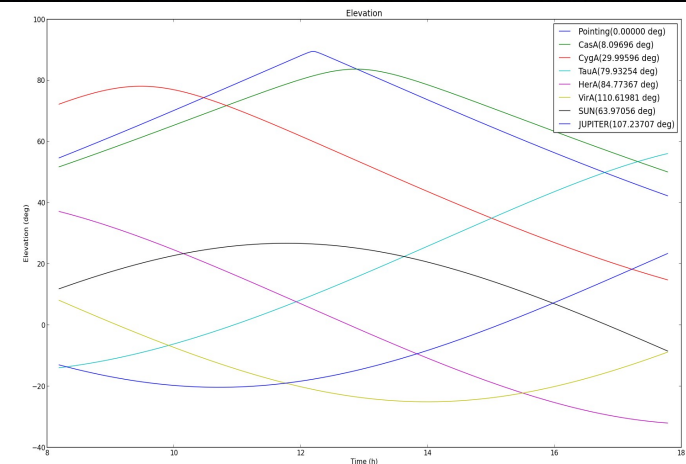
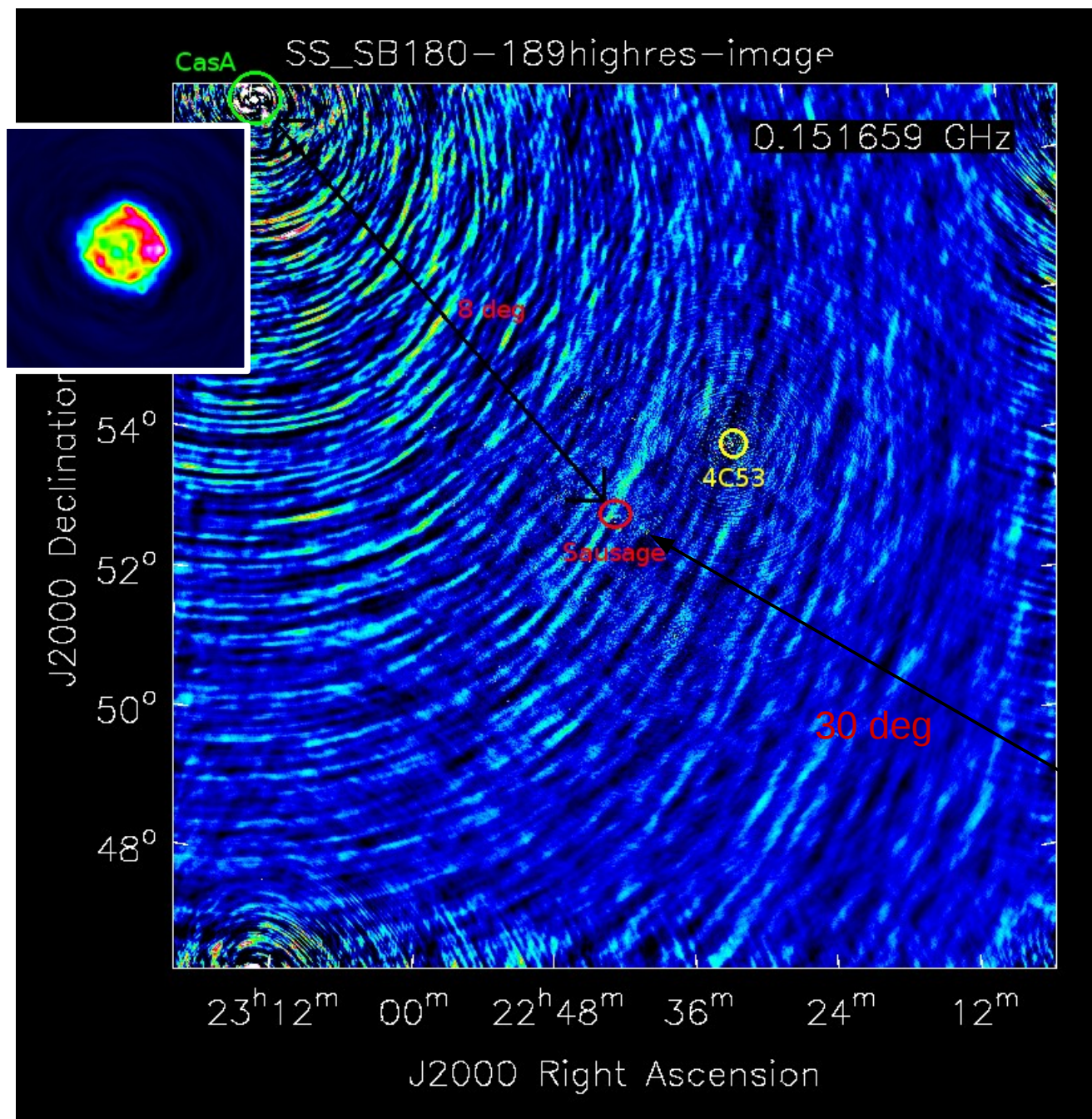
At the shock: particle
acceleration
with a bit of magnetic field
this gives synchrotron
radiation

classification:
relic radio sources

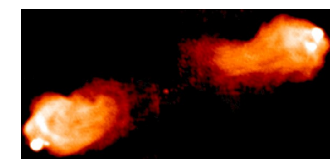


~ 20 years
earlier....(Ger
de Bruyn)



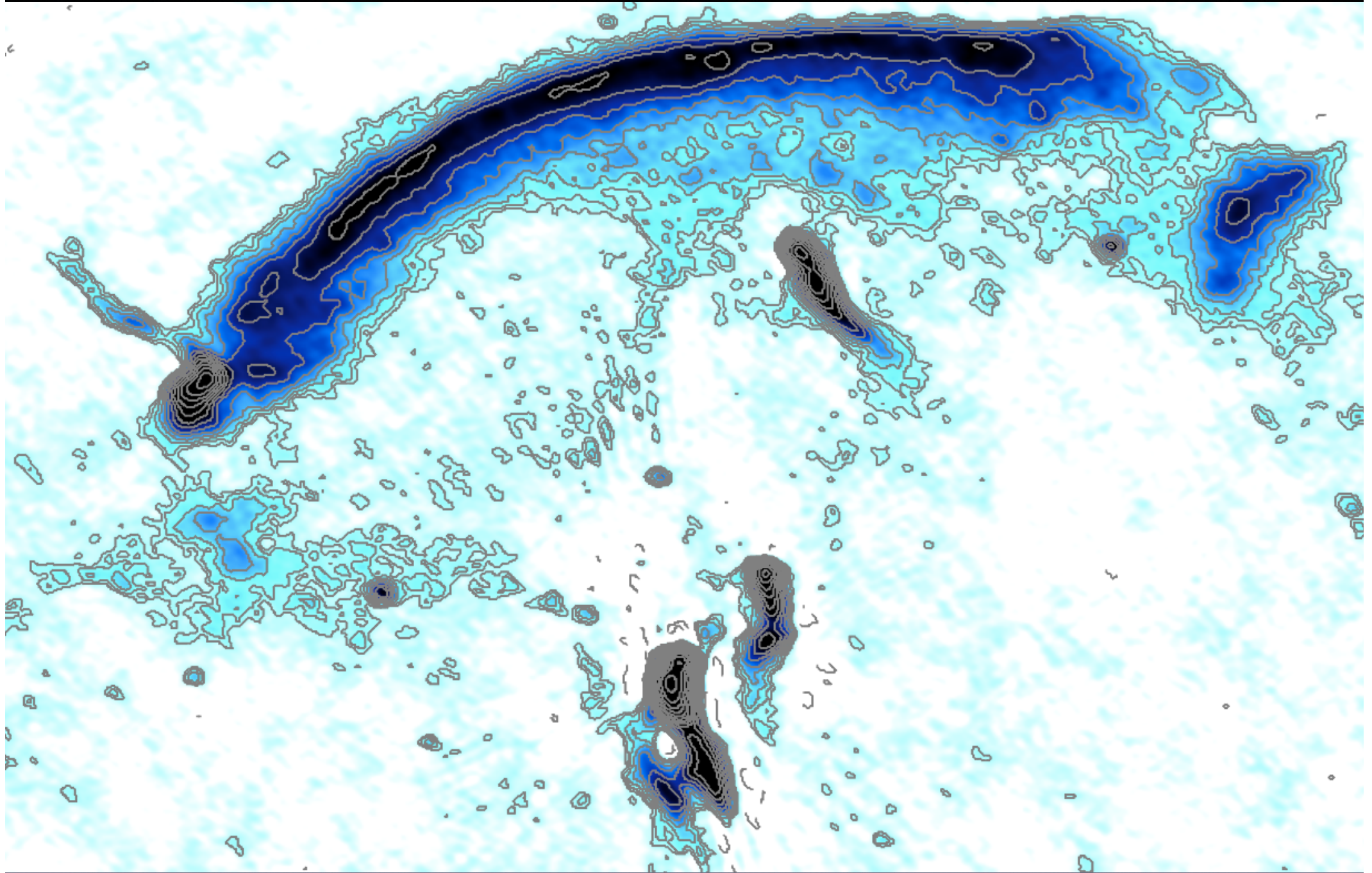


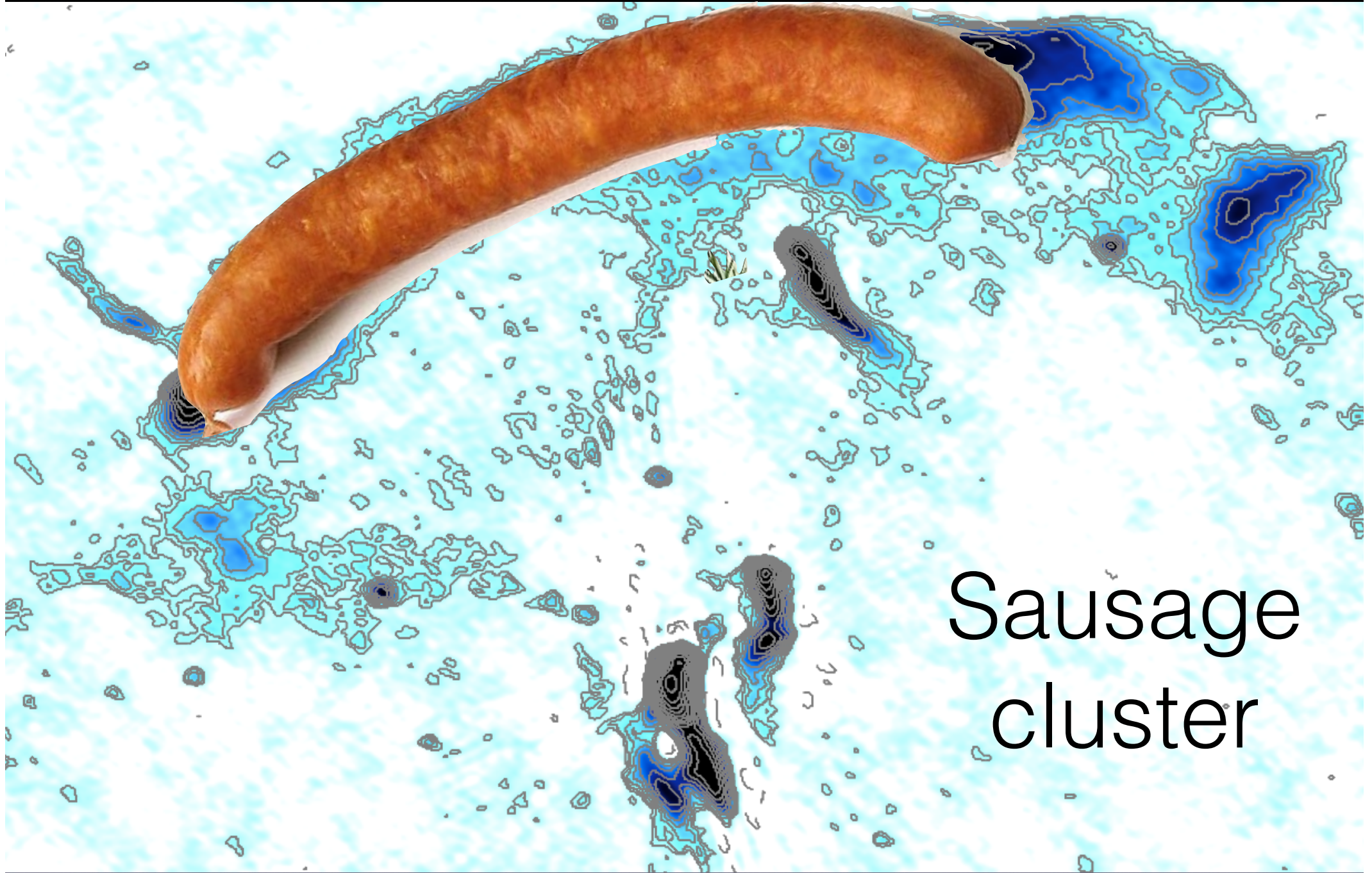
- CasA skymodel: 69MHz, 10" (Reinout van Weeren)
- Data resolution: 64ch, 1s
- Time step: 4s
- Freq. step: 16ch



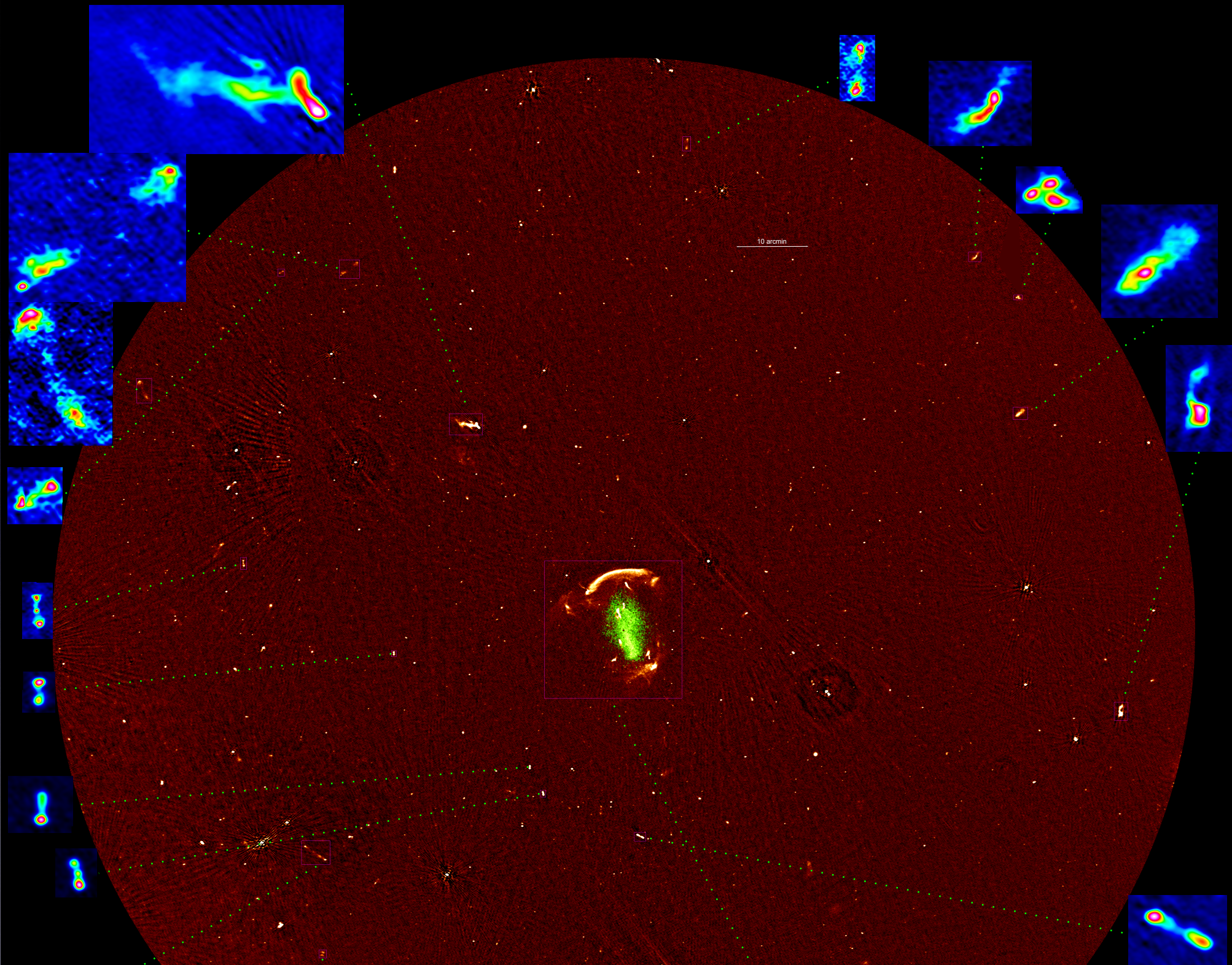
CygA

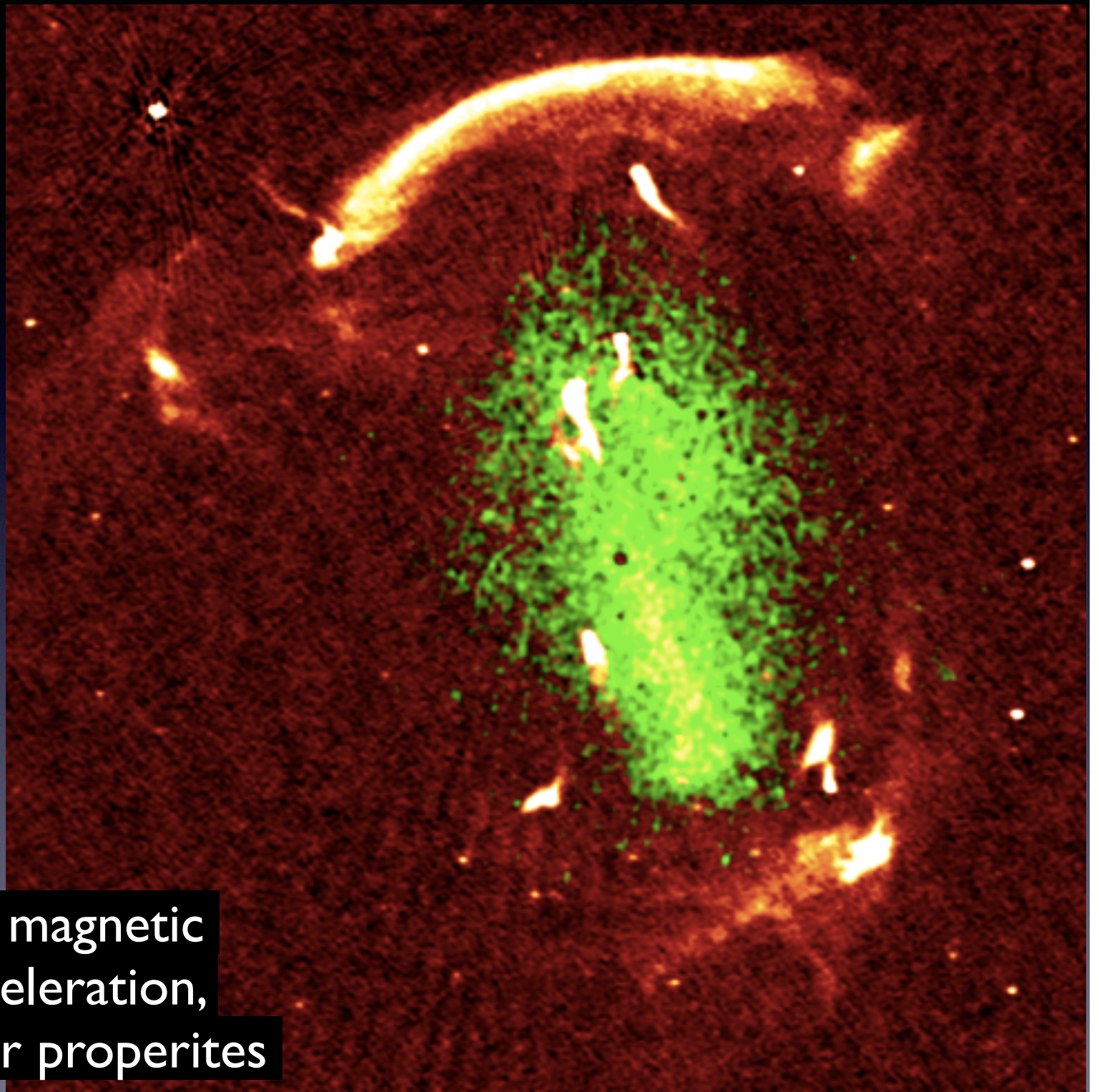






Sausage
cluster





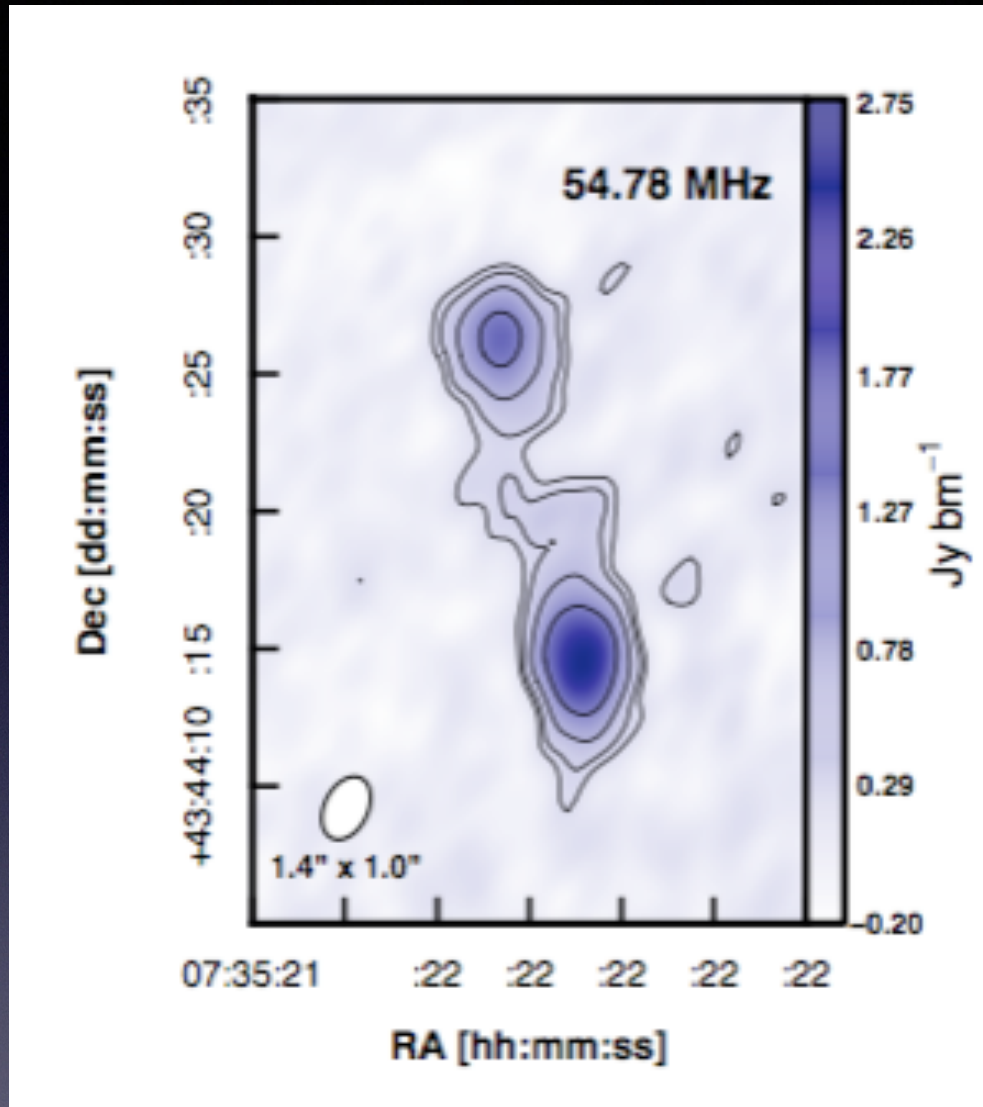
shock properties, magnetic
fields, particle acceleration,
turbulence, cluster properties

Donnert et al. 2016

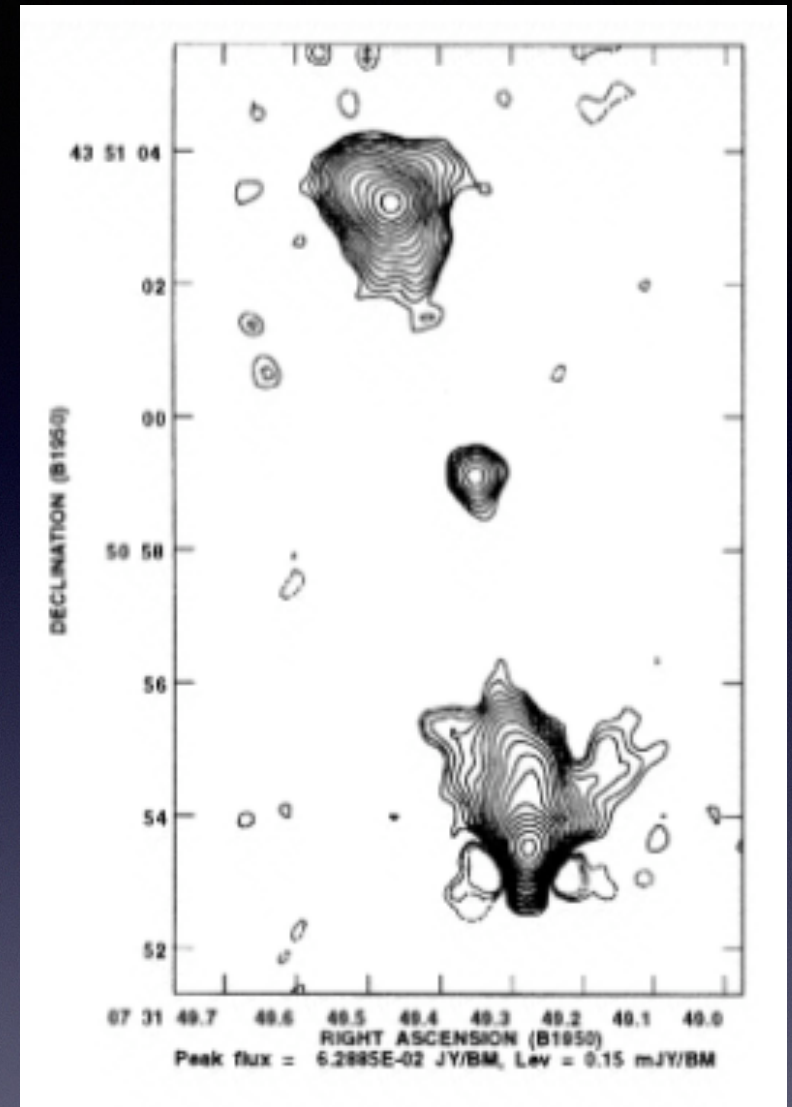
Name	Value	
d_{center}	1.3 Mpc	Distance to center
z	0.188	Redshift
Length	2 Mpc	Longest Extent on the sky
Width	200 kpc	Extent of the brightness profile
Ψ	< 10 deg	Angle into plane of sky
M	4.6	Mach number
T_{up}	3.0 keV	Upstream temperature
$n_{\text{th,up}}$	$1.6 \times 10^{-4} \text{ cm}^{-3}$	Upstream thermal number density
s	2.1	Injection spectral index of CRE
v_{dw}	1184 km/s	Downwind speed
$c_{\text{s,up}}$	902 km/s	Sound speed of upstream medium
v_{shock}	4144 km/s	Shock speed in upstream medium
σ	3.5	Compression ratio
σ_{Temp}	7.5	Thermal compression ratio
$n_{\text{th,dw}}$	$5.6 \times 10^{-4} \text{ cm}^{-3}$	Downwind thermal number density
$T_{\text{th,dw}}$	22.4 keV	Downwind temperature
B_{IC}	4.6 μG	B_{CMB} at $z=0.188$

Plus merger geometry (2:1 mass ratio, zero impact parameter and time since core passage ~ 1 Gyr)

4C+43.15: a powerful radio galaxy at $z=2.5$

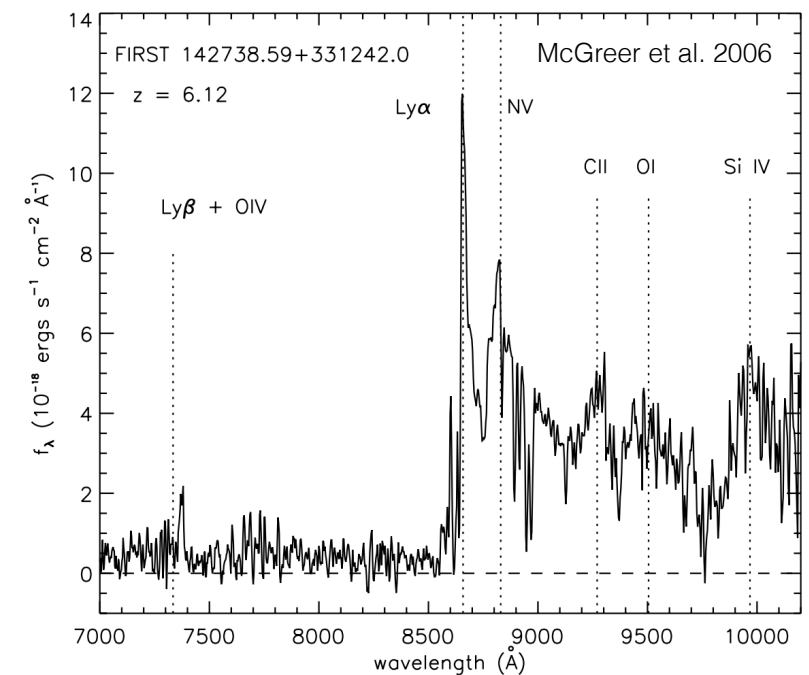
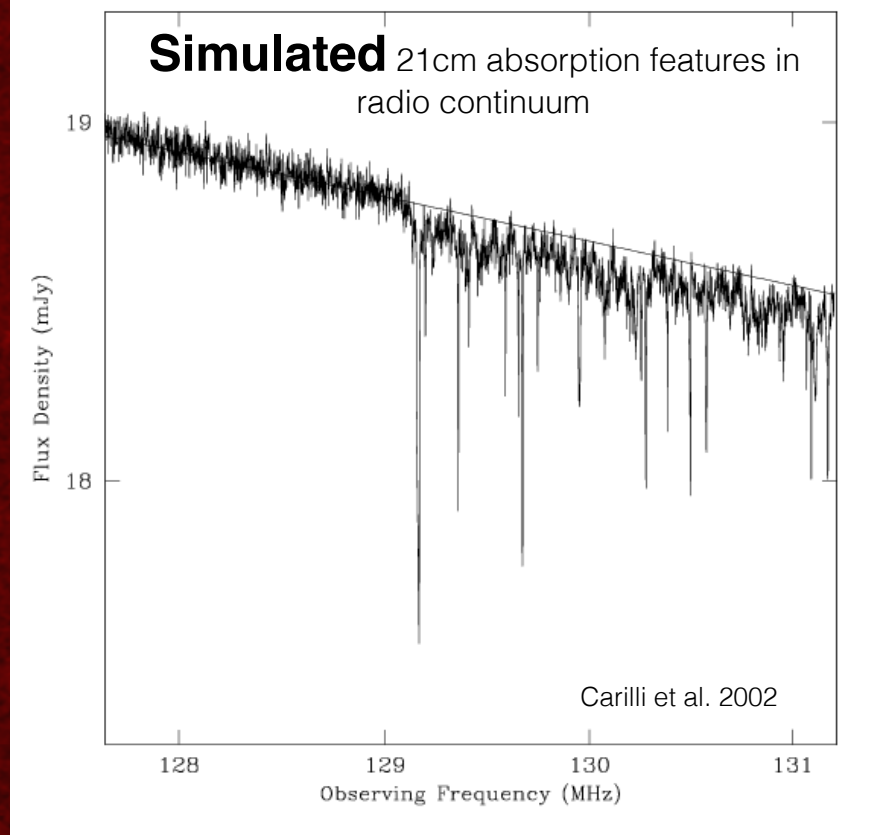


LOFAR 58 MHz, 40 mJy
(2 times thermal noise, 4.5 hrs obs.)
Resolution: 0.9×0.6 arcsec
Morabito et al. in prep



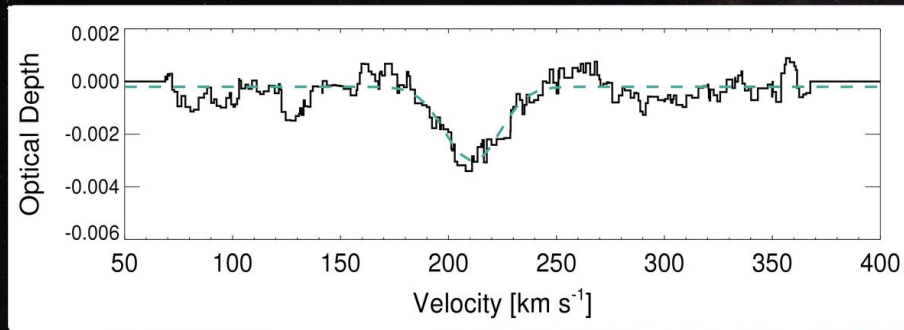
VLA 4.6 GHz
Carilli et al. 1997

$z = 6.12$ Radio Loud QSO as seen in Bootes field

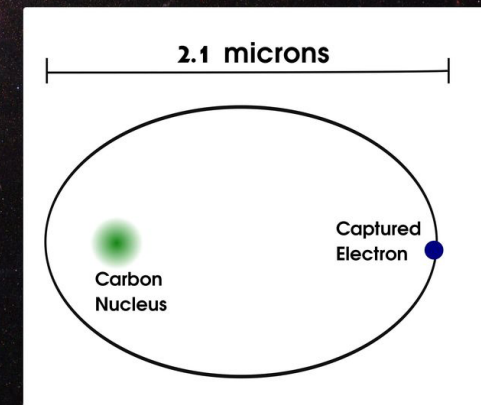


Williams, Retana, Saxena, Duncan

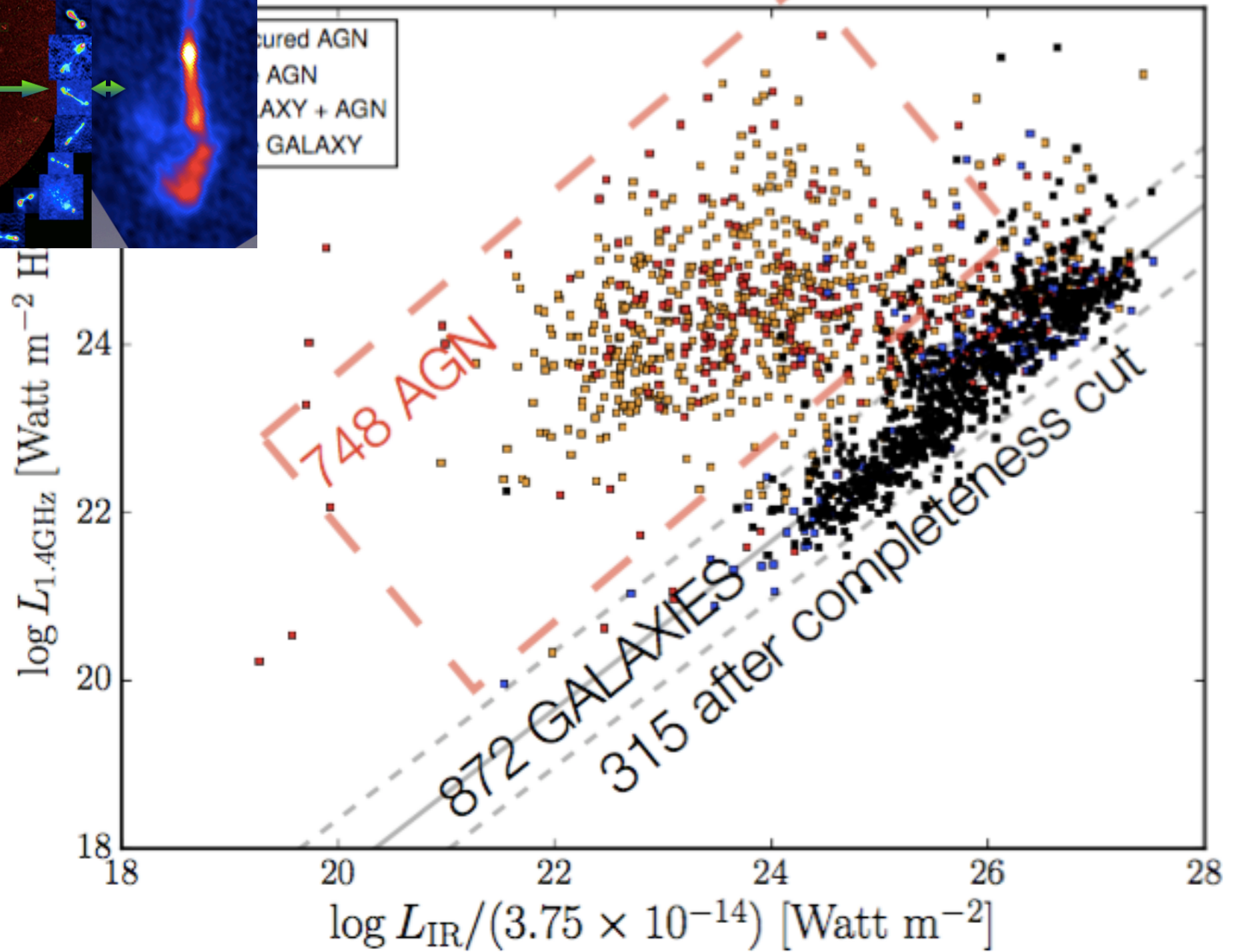
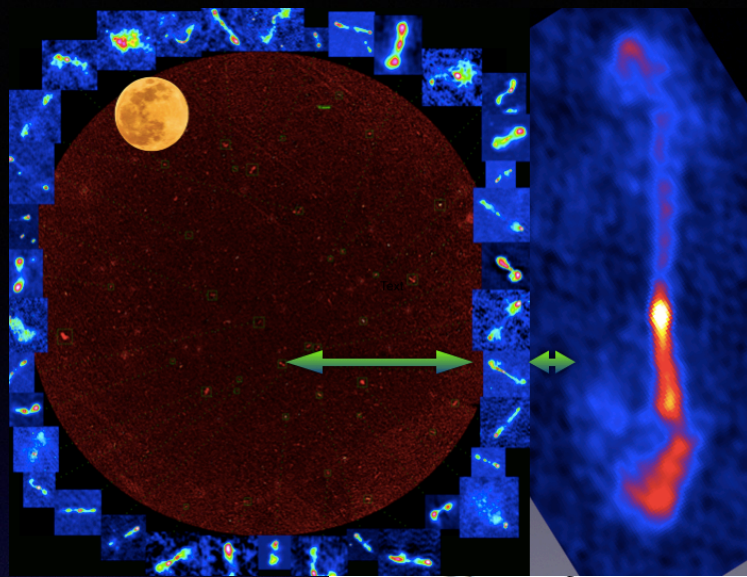
LOFAR discovers largest carbon atoms outside our Milky Way



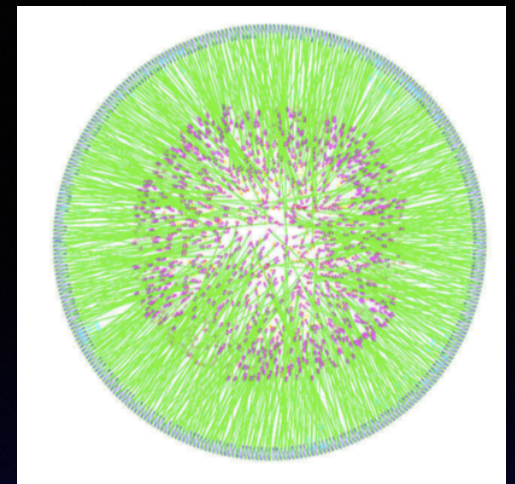
Carbon radio recombination lines ($n \sim 400-500$):
tracing the cold (~ 100 K), diffuse ($n < 1 \text{ cm}^{-3}$) medium



AGN/starburst studies



WEAVE 1000 fibre spectrograph on WHT



- 1 million fibre projects for LOFAR follow-up starting in 2018
- Science
 - Star-formation evolution across cosmic history
 - Co-evolution of black-hole and star formation - physics of AGN feedback
 - Finding radio galaxies in the Epoch of Reionisation at $z > 6$
- Three tier survey design

Deep
(~10s of deg^2)

Mid (650+ deg^2)

Wide (10,000 deg^2)

Euclid map 15,000 degrees^2 in o and IR
for weak lensing and BAOs
Launch 2020